AstroSat Calibration Meeting (AstroSat Science Support Cell, IUCAA, Pune)

Cross spectral calibration of LAXPC, CZTI, and NuSTAR with Crab

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Analysis:

3 epochs of Crab observations with AstroSat are analysed from 2017, 2018, and 2020. Quasi-contemporaneous NuSTAR observations are considered.

LAXPC:

- Data reduced using Format (A) pipeline.
- Only LAXPC20 spectra are used for cross-calibration fitting.

CZTI:

• Spectra produced using pipeline v3 and associated CALDB

NuSTAR:

- nupipeline v0.4.9 (2021-03-21)
- For bright sources, post-processing filters can veto some good source counts. Can underestimate source flux and lead to mismatch between FPMs.

```
statusexpr="STATUS==b0000xxx00xxx000"
```

SXT:

• Pipeline v1.4b

LAXPC-CZTI

- LAXPC20 and CZTI_Q0 were jointly fitted.
- LAXPC (4 50 keV). CZTI (30 100 keV)
- Model: TBabs*powerlaw
- nH fixed at $2.2 \times 10^{21} \text{ cm}^{-2}$ (Madsen et al 2015)
- The photon index and its norm were tied across the instruments





- The cross-normalization constant has increased from 1% to 15% from 2017 to 2020
- The gaussian at 30 keV was not required for 0964.
- Photon index and normalization have remained consistent
- Total (4 100 keV) flux is 3 x 10⁻⁸ erg s⁻¹ cm⁻²

Date	OBSID	Constant	Γ	PL norm	LineE	Sigma	norm	$\chi^2_{ u}$
20170114	0964	0.99 ± 0.01	2.08 ± 0.01	7.39 ± 0.15	-	-	-1	1.01
20180914	2638	1.05 ± 0.03	2.10 ± 0.01	7.57 ± 0.24	30.1 ± 2.4	$7.6^{+3.5}_{-2.6}$	0.012 ± 0.006	1.04
20200913	3866	1.14 ± 0.03	2.10 ± 0.01	7.29 ± 0.22	31.3 ± 1.7	$10.0^{+3.0}_{2.0}$	0.025 ± 0.006	1.04

LAXPC-NuSTAR / CZTI-NuSTAR

- No exact simultaneous data. Joint fitting with nearest NuSTAR observations
- The fitting results are dominated by NuSTAR due to better statistics
- The cross-normalization constant for CZTI remains stable at ~17%. For LAXPC it changes from 19% to 26%.



Energy (keV)

Cross-normalization constant of LAXPC and CZTI with NuSTAR

Instrument	2017	2018	2020
LAXPC	0.813 ± 0.005	0.784 ± 0.006	0.741 ± 0.006
CZTI	0.834 ± 0.004	0.831 ± 0.005	0.838 ± 0.005

LAXPC - NuSTAR

Separately fitted in the common range of 4-55 keV.

- Difference in flux is about 20%
- Fluxes in both LAXPC and NuSTAR decreased from 2017 to 2018/2020.
- Best fit photon index remains stable
- LAXPC contours continually increase in size.





CZTI - NuSTAR

Separately fitted in the range of 30 -78 keV.

- Difference in photon index have reduced, due to decrease in NuSTAR estimates.
- Difference in flux remains at ~15%





Cross-calibration of the four quadrants of CZTI.



Spectral state dependence of cross-normalization factor between SXT and LAXPC A test with Cygnus X-1

- SXT analysis using pipeline v1.4b
- Pileup correction done by using annular region selection.
- Hard state requires 2' 3' inner radius. Softer states require bigger inner radius (~6').



• During hard state the constant factor is about 60%. It is independent of the model and spectral binning.



	Data	group: 1			
constant	factor		1.62788		
TBpcf	nH	10^22	2.52480		
TBpcf	pcf		0.500000		
TBpcf	redshift		0.0		
TBabs	nH	10^22	0.413357		
simpl	Gamma		1.63246		
simpl	FracSctr		0.350431		
simpl	UpScOnly		1.00000		
diskbb	Tin	keV	0.336923		
diskbb	norm		2.47178E+04		
Data group: 2					
constant	factor		1.00000		
TBpcf	nH	10^22	2.52480		
TBpcf	pcf		0.500000		
TBpcf	redshift		0.0		
TBabs	nH	10^22	0.413357		
simpl	Gamma		1.63246		
simpl	FracSctr		0.350431		
simpl	UpScOnly		1.00000		
diskbb	Tin	keV	0.336923		
diskbb	norm		2.47178E+04		
	constant TBpcf TBpcf TBabs simpl simpl diskbb diskbb diskbb constant TBpcf TBpcf TBpcf TBpcf TBpcf TBabs simpl simpl diskbb diskbb	Data constant factor TBpcf nH TBpcf pcf TBpcf redshift TBabs nH simpl Gamma simpl FracSctr simpl UpScOnly diskbb Tin diskbb norm Data constant factor TBpcf nH TBpcf pcf TBpcf redshift TBabs nH simpl Gamma simpl FracSctr simpl UpScOnly diskbb Tin diskbb Tin	Data group: 1 constant factor TBpcf nH 10^22 TBpcf pcf TBpcf redshift TBabs nH 10^22 simpl Gamma simpl FracSctr simpl UpScOnly diskbb Tin keV diskbb norm Data group: 2 constant factor TBpcf nH 10^22 TBpcf pcf TBpcf redshift TBabs nH 10^22 simpl Gamma simpl FracSctr simpl UpScOnly diskbb Tin keV diskbb Tin keV		

- For intermediate and soft states (where disk emission is strong), the cross-normalization factor is 10% 20%
- Gain offset lied between 35-45 eV for all observations.



		Data g	roup: 1			
1	constant	factor	17	1.15575		
2	TBabs	nH	10^22	0.489270		
3	diskbb	Tin	keV	0.284617		
3	diskbb	norm		4.33404E+04		
4	nthComp	Gamma		1.67609		
4	nthComp	kT e	keV	1000.00		
4	nthComp	kT bb	keV	0.284617		
4	nthComp	inp type	0/1	1.00000		
4	nthComp	Redshift		0.0		
4	nthComp	norm		0.828516		
5	relxillCp	Index1		3.00000		
5	relxillCp	Index2		3.00000		
5	relxillCp	Rbr		100.000		
5	relxillCp	a		0.990000		
5	relxillCp	Incl	deg	30.0000		
5	relxillCp	Rin		-2.10216		
5	relxillCp	Rout		400.000		
5	relxillCp	Z		0.0		
5	relxillCp	gamma		1.67609		
5	relxillCp	logxi		4.03481		
5	relxillCp	Afe		10.0000		
5	relxillCp	kTe	keV	1000.00		
5	relxillCp	refl_frac		-1.00000		
5	relxillCp	norm		1.38312E-02		
Data group: 2						
1	constant	factor		1.00000		
2	TBabs	nH	10^22	0.489270		
3	diskbb	Tin	keV	0.284617		
3	diskbb	norm		4.33404E+04		
4	nthComp	Gamma		1.67609		
4	nthComp	kT_e	keV	1000.00		
4	nthComp	kT_bb	keV	0.284617		